

Towards the Development of a Unified Distributed Data System for L1 Spacecraft

Year One Progress Report

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1. Summary and review of objectives

The purpose of this grant, "Towards the Development of a Unified Distributed Data System for L1 Spacecraft", is to take the initial steps towards the development of a data distribution mechanism for making in-situ measurements more easily accessible to the scientific community. Our obligations as subcontractors to this grant are to add our Faraday Cup plasma data to this initial study and to contribute to the design of a general data distribution system.

The year 1 objectives of the overall project as stated in the GSFC proposal are:

1. Both the *rsync* and *Perl* based data exchange tools will be fully developed and tested in our mixed, Unix, VMS, Windows and Mac OS X data service environment. Based on the performance comparisons, one will be selected and fully deployed. Continuous data exchange between all L1 solar wind monitors initiated.
2. Data version metadata will be agreed upon, fully documented, and deployed on our data sites.
3. The first version of the data description rules, encoded in a XML Schema, will be finalized.
4. Preliminary set of library routines will be collected, documentation standards and formats agreed on, and desirable routines that have not been implemented identified and assigned.
5. ViSBARD test site implemented to independently validate data mirroring procedures.

The specific MIT tasks over the duration of this project are the following:

- a) Implement mirroring service for WIND plasma data
- b) Participate in XML Schema development
- c) Contribute toward routine library

2. Description of year 1 work at MIT

Work at MIT in the first year of this project centered on overall project goals 1-3 through work on MIT tasks (a) and (b). The two main components of this work are the creation of an ASCII format text data file for the processed Wind SWE Faraday Cup hydrogen and helium datasets and the installation of the *rsync* data exchange software on the CSR computer network.

2.1 Generation of ASCII data files

Wind/SWE Faraday Cup ion data are converted from raw binary data into structured IDL save files with calibrated measured currents and spacecraft ephemeris. The key parameter "KP"

analysis code (which analyzes the data using a non-linear least-squares fit to an isotropic Maxwellian velocity distribution) run by both MIT and NASA is applied within hours of receiving data from Wind and the resulting KP data products (hydrogen bulk speed, number density, temperature) are available within one day on the web. MIT also produces a more detailed analysis of the same raw data, currently with a delay of several months, in which hydrogen and helium are characterized through non-linear fitting of *bi*-Maxwellian velocity distribution functions. Since 2001 the bi-Maxwellian data have been made available through the NSSDC in yearly increments with delays of at least one year from the arrival of the raw data at MIT.

The data prepared for the NSSDC are in 20-day ASCII files generated by the IDL program *prepare_bimax_nssdc.pro* at the end of each year and in the event of a re-processing of archival data. Under the grant from GSFC, a new program was created, based on the NSSDC data product, called *gen_bimax_ascii_interval.pro*, to generate ASCII data files for individual days. This program has been run on the available data from the start of the Wind mission through summer 2004 and text data files have been generated. In the future we may consider generating smaller, more clearly documented data files using the CDF format, for example. In the mean time we're sticking with the ASCII files and including a header in each file to explain the contents (see Appendix A for an example).

2.2 Installation of rsync server

We have installed the rsync software as requested by Dr. Szabo on a computer on the CSR network, birkeland.mit.edu, with access to the Wind data files. A file system was created “/nfs/faraday/d9/wind/vho/” to house data related to the VHO project. Currently the sole subdirectory of this system is “wind_swe_fc_apbimax”, which holds the Wind SWE data.

We have tested synchronizing data using rsync over the local network, and we have just learned at the Fall 2004 AGU meeting that GSFC was able to connect to MIT and view the available files. Currently access to the data is limited to MIT and the GSFC computers lepszsl.gsfc.nasa.gov and zathras.gsfc.nasa.gov.

3. Plans for future work

We will progress with our future assignments as described in the original grant. In particular we will continue working on an xml description of the SWE plasma data (and plasma data in general). In addition to the xml we have identified the following path to making full use of the rsync system:

1. Implement regular downloading of 3-second MFI data from GSFC.
2. Run bi-Maxwellian analysis of SWE Faraday Cup data as soon as they are available from GSFC as determined by rsync
3. Immediately make processed SWE data available through rsync.

Appendix A: Sample data file header

```
; Project: WIND/SWE Faraday Cup (Proton-Alpha Anisotropy Analysis)
; Description: Solar wind proton and alpha parameters, including anisotropic
; temperatures, derived by non-linear fitting of the measurements
; and with moment techniques.
; Filename:
/nfs/faraday/d9/wind/vho/wind_swe_fc_apbimax/wi_swe_fc_apbimax.2000363.txt
; Author: J. Kasper (jck@mit.edu) 1 (617) 253-7611
```

Created: Fri Dec 10 11:03:42 2004

Notes:

- Data reported within this file do not exceed the limits of various parameters listed in the following section. There may be more valid data in the original dataset that requires additional work to interpret but was discarded due to the limits. In particular we have tried to exclude non-solar wind data from these files.
- We provide the one sigma uncertainty for each parameter produced by the non-linear curve fitting analysis either directly from the fitting or by propagating uncertainties for bulk speeds, flow angles or any other derived parameter.
- For the non-linear anisotropic proton analysis, a scalar thermal speed is produced by determining parallel and perpendicular temperatures, taking the trace, $T_{\text{scalar}} = (2T_{\text{perp}} + T_{\text{para}})/3$ and converting the result back to a thermal speed. The uncertainties are also propagated through

Limits:

Minimum mach number: 0.75000000
Maximum chisq/dof: 100000.00
Minimum distance
to bow shock: -100000.00 [Re]
Maximum uncertainty in any
parameter from non-linear
analysis: 70.0000[%]

Total number of spectra in this year: 864
Spectra after applying limits: 864
Spectra removed as time regressions: 0
Spectra removed as spikes: 11
Spectra written to this file: 847
Spectra with alpha particle data: 418

Format Code: (i4.4," ",f11.6,18(" ",f7.1),2(" ",f9.3),10(" ",f7.1),2(" ",f9.3)," ",f6.1,7(" ",f7.1)," ",f9.3,10(" ",f9.3))

Column # Description

1	Year
2	Fractional day of year of start of spectrum (Noon on January 1 = 1.5)
3	Proton bulk speed V (km/s) from non-linear analysis
4	One sigma uncertainty in V [km/s]
5	Proton velocity component Vx (GSE, km/s) from non-linear analysis
6	One sigma uncertainty in Vx [km/s]
7	Proton velocity component Vy (GSE, km/s) from non-linear analysis
8	One sigma uncertainty in Vy [km/s]
9	Proton velocity component Vz (GSE, km/s) from non-linear analysis
10	One sigma uncertainty in Vz [km/s]
11	Scalar proton thermal speed (km/s) trace of anisotropic temperatures

12	One sigma uncertainty in trace thermal speed
13	Proton thermal speed $W_{\text{perpendicular}}$ (km/s) from non-linear analysis
14	One sigma uncertainty in $W_{\text{perpendicular}}$ [km/s]
15	Proton thermal speed W_{parallel} (km/s) from non-linear analysis
16	One sigma uncertainty in W_{parallel} [km/s]
17	East-West flow angle (degrees)
18	One sigma uncertainty in EW flow angle
19	North-South flow angle (degrees)
20	One sigma uncertainty in NS flow angle
21	Proton number density N_p (n/cc) from non-linear analysis
22	One sigma uncertainty in N_p [n/cc]
23	Alpha bulk speed V (km/s) from non-linear analysis
24	One sigma uncertainty in V [km/s]
25	Alpha velocity component V_x (GSE, km/s) from non-linear analysis
26	One sigma uncertainty in V_x [km/s]
27	Alpha velocity component V_y (GSE, km/s) from non-linear analysis
28	One sigma uncertainty in V_y [km/s]
29	Alpha velocity component V_z (GSE, km/s) from non-linear analysis
30	One sigma uncertainty in V_z [km/s]
31	Scalar alpha thermal speed (km/s) trace of anisotropic temperatures
32	One sigma uncertainty in trace thermal speed
33	Alpha number density N_a (n/cc) from non-linear analysis
34	One sigma uncertainty in N_a [n/cc]
35	CHISQ/DOF for this fit
36	Proton bulk speed (km/s) from moment analysis
37	Proton velocity component V_x (GSE, km/s) from moment analysis
38	Proton velocity component V_y (GSE, km/s) from moment analysis
39	Proton velocity component V_z (GSE, km/s) from moment analysis
40	Proton thermal speed W (km/s) from isotropic moment analysis
41	Proton thermal speed $W_{\text{perpendicular}}$ (km/s) from bimax moment analysis
42	Proton thermal speed W_{parallel} (km/s) from bimax moment analysis
43	Proton number density N_p (n/cc) from moment analysis
44	Magnetic field component B_x (GSE, nT) averaged over plasma measurement
45	Magnetic field component B_y (GSE, nT) averaged over plasma measurement
46	Magnetic field component B_z (GSE, nT) averaged over plasma measurement
47	Angular deviation of magnetic field over plasma measurement [degrees]
48	Deviation in magnitude of field over plasma measurement [nT]
49	X (GSE) Position of Wind S/C at start of spectrum [Re]
50	Y (GSE) Position of Wind S/C at start of spectrum [Re]
51	Z (GSE) Position of Wind S/C at start of spectrum [Re]
52	Y (GSM) Position of Wind S/C at start of spectrum [Re]
53	Z (GSM) Position of Wind S/C at start of spectrum [Re]